

Strong Noise – Removal and Replacement on Seismic Data

Patrick Butler, GEDCO, Calgary, Alberta, Canada

pbutler@gedco.com

Summary

A module for removing and replacing strong noise in seismic data is presented. It uses thresholded median replacement in the frequency domain (Bekara et al., 2007 and Elboth et al., 2010). By using short FFT windows in time we both combine the samples belonging to a single event and insure that if the amplitudes are modified the output will still be smooth in time. If the median with a threshold for replacement is used to remove strong noise on CDP gathers, stack, prestack migration, white noise suppression can be used to suppress the remaining white noise. Because the module does not depend on noise coherency, removal of offline energy, noise bursts, ground roll, and repeats of first breaks is very effective. Replacing all the data with the median does work but the result is somewhat synthetic.

In short this methodology is robust and effective and is recommended for routine use in the processing of seismic data.

Introduction

Strong noise can overwhelm stack, prestack migration and other linear processes. The root n cancellation for Gaussian noise may just not be sufficient. To address this we created a nonlinear process named THOR.

Theory and/or Method

The process must be nonlinear - the noise is such that linear processes like stack and migration are failing. It is not wanted to impact traces from which signal can be recovered so a threshold is needed above which the trace is recognized as badly contaminated. Data could be edited and to preserve normalization stack, migration could remember the trace weighting but this process does not use that approach. Instead it progressively replaces the trace data with an estimate of the signal built from adjacent traces.

If the data has NMO applied then, if it is sorted into CDP mode, the signal will be relatively consistent across a time slice. In Figure 1 a CDP gather with noise bursts that have been replaced is shown.

A median could be used but medians rattle about as the basic noise on the trace. Being applied successively at each sample in time will produce a high frequency rattle in time as is shown in Figure 2. Notice that the THOR replacement spectrum is much closer to that of the stack than the median. The stack differs from the THOR result mostly because THOR is looking at only 7 traces as opposed to the stack of 34 traces.

Noise Burst Replacement

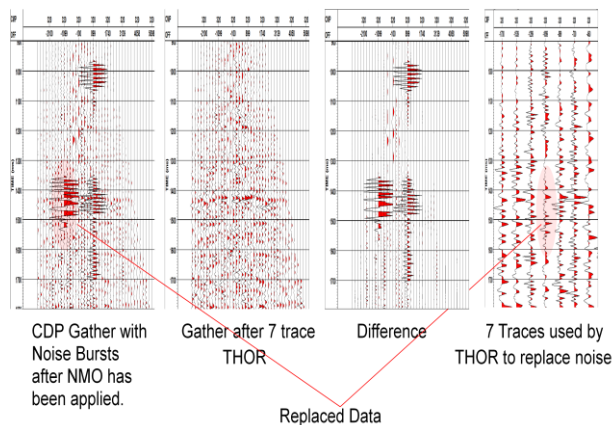


Figure 1: a CDP gather with noise bursts

Replace Single Noise Burst

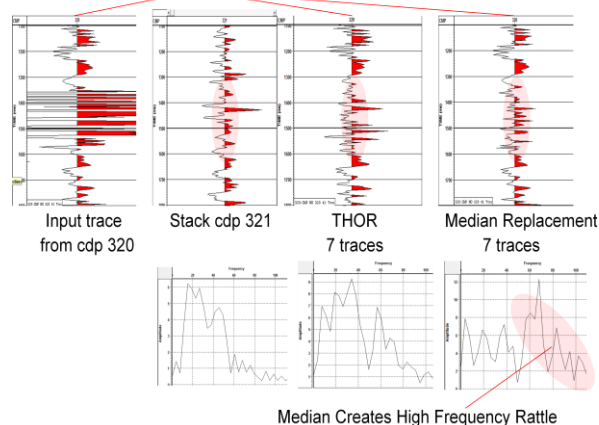


Figure 2: THOR vs. time domain median

To show what the THOR process actually produces the threshold has been set to near to zero and all the data in the gather has been replaced. (see Figure 3) Note that the process is following the signal character in space and hence is AVO friendly with an appropriate median length. The module automatically sorts the data in offset and if longer medians are needed it has the capability to supergather CDPs.

To see how this process works it must be realized that the signal is a wavelet. More than one sample in time must be considered if it is to be recognized and estimated in the presence of background noise. Considering that the wavelet is short and considering that the noise may have restricted frequency content, the data is FFTed in small overlapping windows in time as is shown in Figure 4.

THOR - Replace the Whole Gather

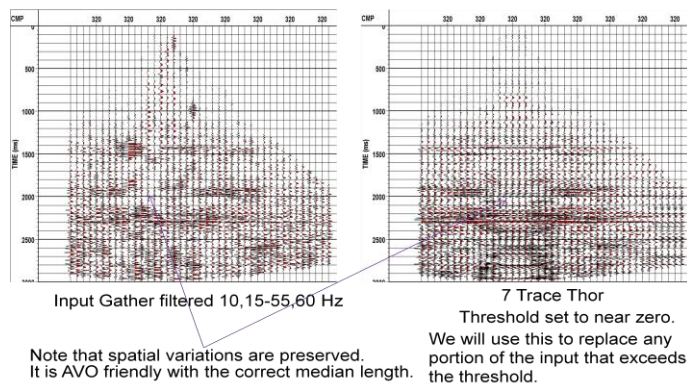


Figure 3: replace all the samples in a gather with THOR

A Single Window

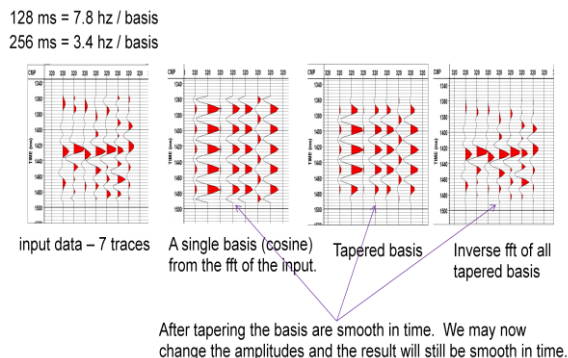


Figure 4: THOR FFTs the input in small overlapping windows

There must be only a small amount of data in each window to insure that the amplitudes for each basis are consistent from trace to trace. A median can now be applied to these amplitudes separately and the rattle that appeared formerly in time has now been transferred into frequency. As you have seen on the noise replacement slides our output in the time domain is now smooth whether we use the original amplitudes or whether we replace some of them with the median amplitudes.

Examples

We now have a good way to run a median across our CDP gathers and replace values if the median minus the trace value exceeds a threshold. Some stacked results are shown in Figure 5.

Stacks without and with THOR

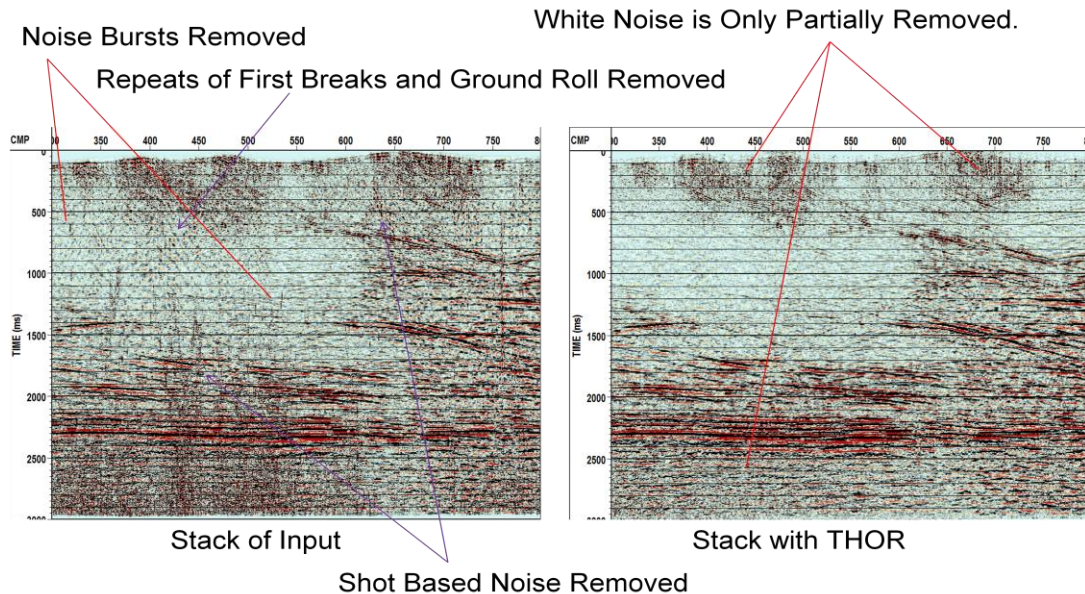


Figure 5: stacks of structured data with and without THOR

As expected the single CDP noise bursts seen previously have been completely removed as has the shot based noise. Note that THOR is not a dip filter. It looks horizontally across CDP gathers. Successive traces in a shot gather fall into different CDPs so the noise is not coherent within any single CDP.

The ground roll and repeats of first breaks have similarly been easily removed. They are restricted in frequency so the median need only attack those frequencies and leave the others untouched. Even if this noise pervades all the shots we can often get a reasonable estimate of the signal underneath it for those frequencies. This occurs because the sine and cosine terms are considered separately. If the noise on a trace is primarily say cosine then the sine term will show the signal amplitude. Provided the noise in CDP mode is incoherent horizontally and provided enough traces are looked at the low frequency signal amplitudes can be estimated.

Notice also that the white noise has only been partially removed. This is deliberate. If the threshold were set low enough to remove it, then it would force the stack to a median solution. As can be seen in Figure 6 this tends to produce a blocky and synthetic looking section. So in general the median is used to remove the strong noise and the average (stack, prestack migration, white noise suppression) is used to deal with the white noise problems.

Figure 7 shows that THOR makes a sensible estimate of the signal low frequencies underneath the ground roll.

THOR Stack vs full Median

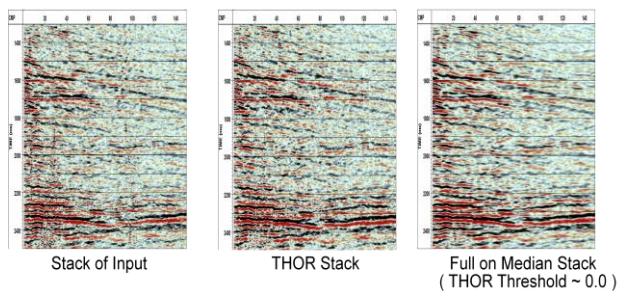


Figure 6: removing just the strong noise is a bit better

Shot Noise

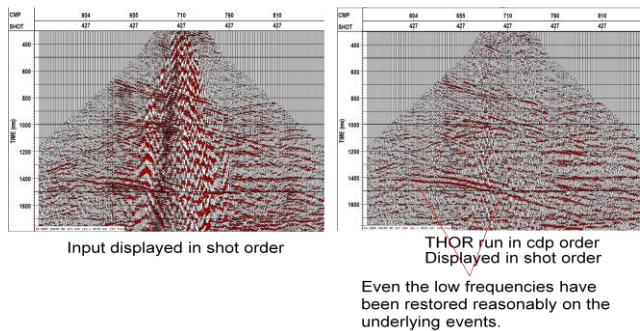


Figure 7: attacking ground roll

Figure 8 shows an offline secondary source that is being excited by every shot on the line. Because the noise is not consistent across the CDP gathers the apex of the noise is being removed just as easily as the diffraction wings.

In Figure 9 the data looks very bad in a CDP gather and THOR has made a substantial difference. However for the stack, the gain is not as good as might be expected mostly because high fold does cancel noise. Low fold at the edges is also causing problems.

Shot Noise - Secondary Source

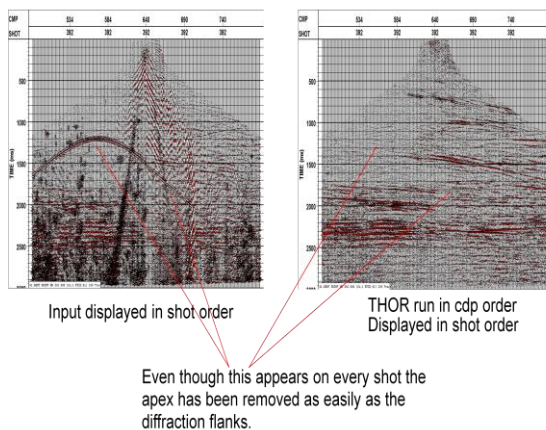


Figure 8: removal of offline energy

Strong Noise on a High Fold 3D

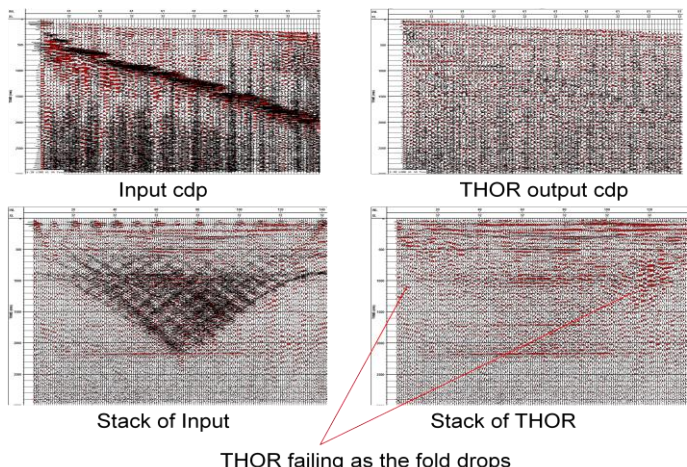


Figure 9: stack is a good noise suppressor on high fold data

Conclusions

By using short FFT windows in time both the samples belonging to a single event are combined and if the amplitudes are modified the output will still be smooth in time. With this decomposition a median combined with a threshold can be used to replace strong noise in CDP gathers. Because the method does not depend on noise coherency removal of offline energy, noise bursts, ground roll and repeats of first breaks is very effective. Replacing all the data with the median does work but is somewhat synthetic.

In short THOR is a robust and effective technology and is suitable for routine processing of seismic data.

Acknowledgements

Thanks to GEDCO for allowing me to present the details of module THOR from the seismic processing package VISTA[®] from GEDCO.

References

Bekara, M., van der Baan, M., 2007, Local singular value decomposition for signal enhancement of seismic data. *Geophysics*, **72**(2), V59-V65.

Elboth, T., Presterud, I.V., Hermansen, D., 2009, Time-frequency seismic data de-noising, *Geophysical Prospecting*, **58**, P.441-453